### Fluid Mechanics - MTF053 Lecture 22

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# Chapter 9 - Compressible Flow



# Learning Outcomes

- 4 Be **able to categorize** a flow and **have knowledge about** how to select applicable methods for the analysis of a specific flow based on category
- 37 **Understand and explain** basic concepts of compressible flows (the gas law, speed of sound, Mach number, isentropic flow with changing area, normal shocks, oblique shocks, Prandtl-Meyer expansion)

Let's go supersonic ...

### Roadmap - Compressible Flow



## **Expansion Waves**

- Gradual change of flow angle
- Increasing flow area
- Increasing Mach number
- Accumulation of infinitesimal flow deflections isentropic

## **Expansion Waves**

What is an expansion wave or expansion region?



# The Prandtl-Meyer Function

- The change of flow properties over an expansion region can be calculated using the Prandtl-Meyer function
- The Prandtl-Meyer function derivation is based on the fact that each expansion wave gives an infinitesimal change in flow angle and flow properties



For small deflection angles, linearization of the  $\theta$ - $\beta$ -Mach relation gives

$$rac{d
ho}{
ho}pprox rac{\gamma \mathcal{M}^2}{(\mathcal{M}^2-1)^{1/2}} d heta$$

The momentum equation for inviscid flows gives

$$dp = -d(\rho V^{2}) = -\rho V dV - V \underbrace{d(\rho V)}_{=0} = -\rho V dV = -\rho V^{2} \frac{dV}{V} = -\rho a^{2} M^{2} \frac{dV}{V} \Rightarrow$$
$$\Rightarrow \{\rho a^{2} = \rho \gamma RT = \gamma p\} \Rightarrow \frac{dp}{p} = -\gamma M^{2} \frac{dV}{V}$$



Now, setting the two expressions for dp/p equal

$$-\gamma M^2 \frac{dV}{V} = \frac{\gamma M^2}{(M^2 - 1)^{1/2}} d\theta \Rightarrow d\theta = -(M^2 - 1)^{1/2} \frac{dV}{V}$$

$$V = Ma \Rightarrow dV = adM + Mda \Rightarrow \frac{dV}{V} = \frac{dM}{M} + \frac{da}{a}$$

$$d\theta = -(M^2 - 1)^{1/2} \left(\frac{dM}{M} + \frac{da}{a}\right)$$



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$$\frac{a_0}{a} = \left(1 + \frac{\gamma - 1}{2}M^2\right)^{1/2}$$

$$da = \left(1 + \frac{\gamma - 1}{2}M^2\right)^{-1/2} da_0 + a_0 d \left[\left(1 + \frac{\gamma - 1}{2}M^2\right)^{-1/2}\right]$$
isentropic  $\Rightarrow da_0 = 0$ 

$$\frac{da}{a} = \frac{d \left[\left(1 + \frac{\gamma - 1}{2}M^2\right)^{-1/2}\right]}{\left(1 + \frac{\gamma - 1}{2}M^2\right)^{-1/2}} = \frac{-\frac{1}{2}\left(1 + \frac{\gamma - 1}{2}M^2\right)^{-3/2}(\gamma - 1)MdM}{\left(1 + \frac{\gamma - 1}{2}M^2\right)^{-1/2}}$$
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$$d\theta = -(M^2 - 1)^{1/2} \left(\frac{dM}{M} + \frac{da}{a}\right)$$

$$\frac{da}{a} = \frac{-\frac{1}{2}(\gamma - 1)MdM}{1 + \frac{\gamma - 1}{2}M^2} \Rightarrow d\theta = -\frac{2(M^2 - 1)^{1/2}}{2 + (\gamma - 1)M^2}\frac{dM}{M}$$

Introducing  $\omega$  defined such that:  $d\omega = -d\theta$ ,  $\omega = 0$  when M = 1

$$\int_{0}^{\omega} d\omega = \int_{1}^{M} \frac{2(M^{2} - 1)^{1/2}}{2 + (\gamma - 1)M^{2}} \frac{dM}{M}$$
$$\omega(M) = \left(\frac{\gamma + 1}{\gamma - 1}\right)^{1/2} \tan^{-1} \left(\frac{M^{2} - 1}{(\gamma + 1)/(\gamma - 1)}\right)^{1/2} - \tan^{-1}(M^{2} - 1)^{1/2}$$
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### The Prandtl-Meyer Function

$$\omega(M) = \left(\frac{\gamma+1}{\gamma-1}\right)^{1/2} \tan^{-1} \left(\frac{M^2-1}{(\gamma+1)/(\gamma-1)}\right)^{1/2} - \tan^{-1}(M^2-1)^{1/2}$$

Prandtl-Meyer function ( $\gamma = 1.4$ )







# Prandtl-Meyer Expansion Waves





# Prandtl-Meyer Expansion Waves

Since the flow is isentropic, the isentropic relations apply:

( $T_o$  and  $p_o$  are constant)

Calorically perfect gas:



$$\frac{T_o}{T} = \left[1 + \frac{1}{2}(\gamma - 1)M^2\right]$$
$$\frac{\rho_o}{\rho} = \left[1 + \frac{1}{2}(\gamma - 1)M^2\right]^{\frac{\gamma}{\gamma - 1}}$$

## Prandtl-Meyer Expansion Waves

since  $T_{o_1} = T_{o_2}$  and  $p_{o_1} = p_{o_2}$ 

$$\frac{T_1}{T_2} = \frac{T_{o_2}}{T_{o_1}} \frac{T_1}{T_2} = \left(\frac{T_{o_2}}{T_2}\right) / \left(\frac{T_{o_1}}{T_1}\right) = \left[\frac{1 + \frac{1}{2}(\gamma - 1)M_2^2}{1 + \frac{1}{2}(\gamma - 1)M_1^2}\right]$$

$$\frac{D_1}{D_2} = \frac{p_{O_2}}{p_{O_1}} \frac{p_1}{p_2} = \left(\frac{p_{O_2}}{p_2}\right) \middle/ \left(\frac{p_{O_1}}{p_1}\right) = \left[\frac{1 + \frac{1}{2}(\gamma - 1)M_2^2}{1 + \frac{1}{2}(\gamma - 1)M_1^2}\right]^{\frac{\gamma}{\gamma - 1}}$$

# Diamond-Wedge Airfoil



**Note!** symmetric airfoil at zero incidence  $\Rightarrow$  zero lift but what about drag?

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- 1-2 standard oblique shock calculation for flow deflection angle  $\varepsilon$  and upstream Mach number  $M_1$
- 2-3 Prandtl-Meyer expansion for flow deflection angle  $2\varepsilon$  and upstream Mach number  $M_2$

3-4 standard oblique shock calculation for flow deflection angle  $\varepsilon$  and upstream Mach number  $M_3$ 

# Diamond-Wedge Airfoil - Wave Drag

Since conditions 2 and 3 are constant in their respective regions, we obtain:

$$D = 2 [p_2 L \sin(\varepsilon) - p_3 L \sin(\varepsilon)] = 2(p_2 - p_3) \frac{t}{2} = (p_2 - p_3)t$$

For supersonic free stream ( $M_1 > 1$ ), with shocks and expansion fans according to figure we will always find that  $\rho_2 > \rho_3$ 

which implies D > 0

Wave drag (drag due to flow loss at compression shocks)

# Flat-Plate Airfoil



It seems that the angle of the flow downstream of the flat plate would be different than the angle of the flow upstream of the plate. Can that really be correct?!



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For the flow in the vicinity of the plate this is the correct picture. Further out from the plate, shock and expansion waves will interact and eventually sort the mismatch of flow angles out

### Flat-Plate Airfoil

▶ Flow states 4 and 5 must satisfy:

- ▶  $p_4 = p_5$
- flow direction 4 equals flow direction 5 ( $\Phi$ )
- Shock between 2 and 4 as well as expansion fan between 3 and 5 will adjust themselves to comply with the requirements

▶ For calculation of lift and drag only states 2 and 3 are needed

States 2 and 3 can be obtained using standard oblique shock formulas and Prandtl-Meyer expansion

# **Oblique Shocks and Expansion Waves**



### Roadmap - Compressible Flow



# Supersonic Stereo

What if you somehow managed to make a stereo travel at twice the speed of sound, would it sound backwards to someone who was just casually sitting somewhere as it flies by? -Tim Currie

Technically, anyway. It would be pretty hard to hear.

Ves.

The basic idea is pretty straightforward. The stereo is going faster than its own sound, so it will reach you first, followed by the sound it emitted one second ago, followed by the sound it emitted two seconds ago, and so forth.



The problem is that the stereo is moving at Mach 2, which means that two seconds ago, it was over a kilometer away. It's hard to hear music from that distance, particularly when your ears were just hit by (a) a sonic boom, and (b) pieces of a rapidly disintegrating stereo.

Wind speeds of Mach 2 would messily disassemble most consumer electronics. The force of the wind on the body of the stereo is roughly comparable to that of a dozen people standing on it:



