

Compressible Flow - TME085

Lecture 2

Niklas Andersson

Chalmers University of Technology
Department of Mechanics and Maritime Sciences
Division of Fluid Mechanics
Gothenburg, Sweden

`niklas.andersson@chalmers.se`

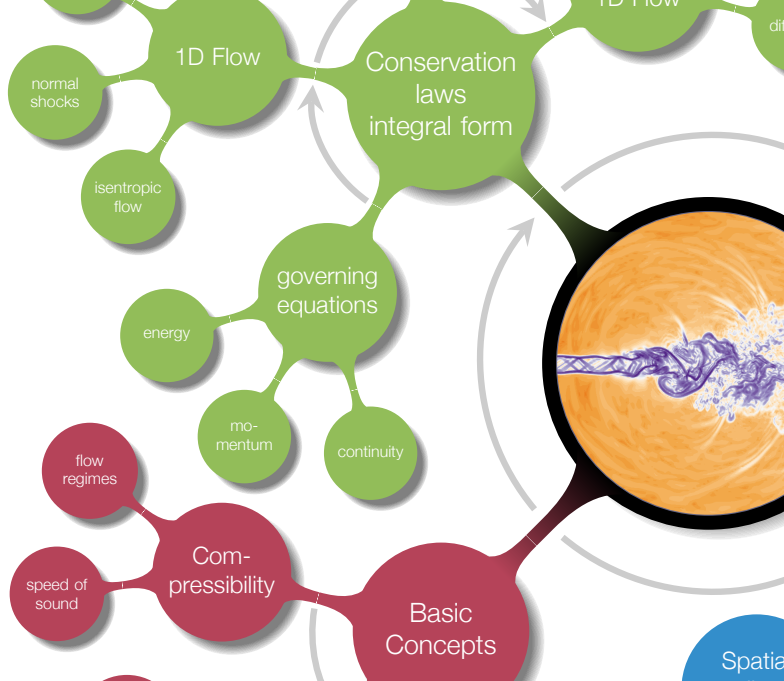


Chapter 2

Integral Forms of the Conservation Equations for Inviscid Flows



Overview

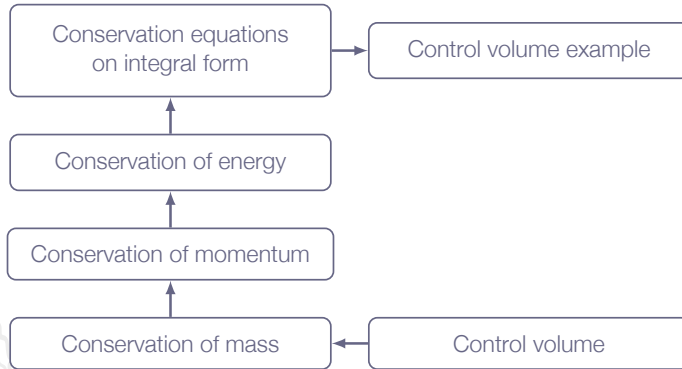


Learning Outcomes

- 4 **Present** at least two different formulations of the governing equations for compressible flows and **explain** what basic conservation principles they are based on
- 5 **Explain** how thermodynamic relations enter into the flow equations
- 7 **Explain** why entropy is important for flow discontinuities

equations, equations and more equations

Roadmap - Integral Relations

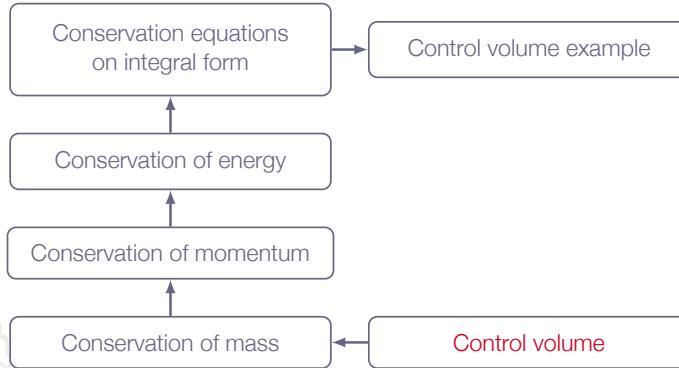


Motivation

We need to formulate the basic form of the governing equations for compressible flow before we get to the applications



Roadmap - Integral Relations



Integral Forms of the Conservation Equations

Conservation principles:

- ▶ conservation of mass
- ▶ conservation of momentum (*Newton's second law*)
- ▶ conservation of energy (*first law of thermodynamics*)



Integral Forms of the Conservation Equations

The control volume approach:

Notation:

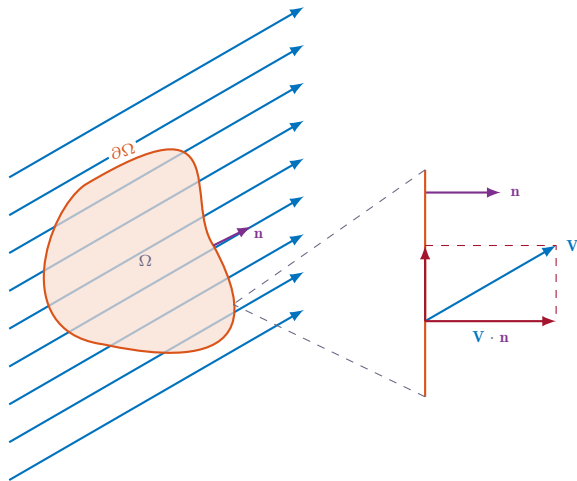
Ω : fixed control volume

$\partial\Omega$: boundary of Ω

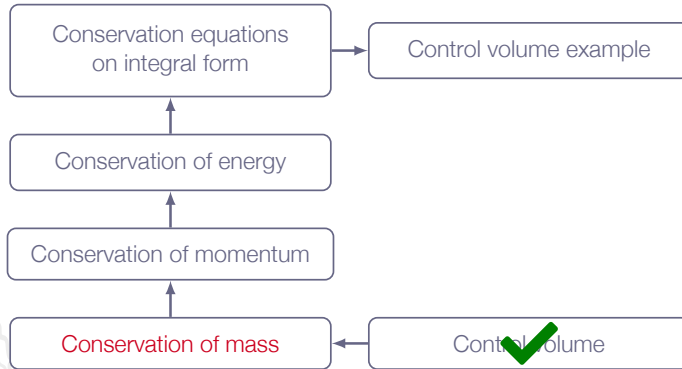
\mathbf{n} : outward facing unit normal vector

\mathbf{v} : fluid velocity

$v = |\mathbf{v}|$



Roadmap - Integral Relations



Chapter 2.3

Continuity Equation



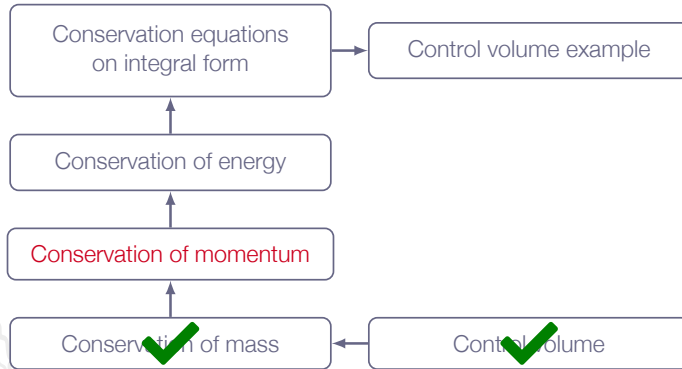
Continuity Equation

Conservation of mass:

$$\underbrace{\frac{d}{dt} \iiint_{\Omega} \rho d\mathcal{V}}_{\text{rate of change of total mass in } \Omega} + \underbrace{\iint_{\partial\Omega} \rho \mathbf{v} \cdot \mathbf{n} dS}_{\text{net mass flow out from } \Omega} = 0$$

Note! notation in the text book $\mathbf{n} \cdot d\mathbf{S} = d\mathbf{S}$

Roadmap - Integral Relations



Chapter 2.4

Momentum Equation



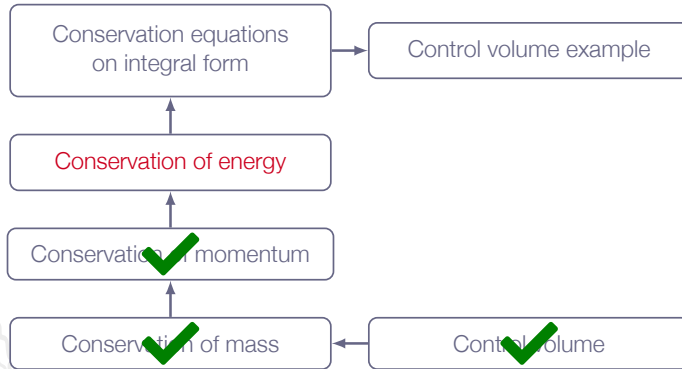
Momentum Equation

Conservation of momentum:

$$\underbrace{\frac{d}{dt} \iiint_{\Omega} \rho \mathbf{v} d\mathcal{V}}_{\text{rate of change of total momentum in } \Omega} + \underbrace{\oint_{\partial\Omega} [\rho(\mathbf{v} \cdot \mathbf{n})\mathbf{v} + p\mathbf{n}] dS}_{\text{net momentum flow out from } \Omega \text{ plus surface force on } \partial\Omega \text{ due to pressure}} = \underbrace{\iiint_{\Omega} \rho \mathbf{f} d\mathcal{V}}_{\text{rate of momentum generation due to forces inside } \Omega}$$

Note! friction forces due to viscosity are not included here. To account for these forces, the term $-(\boldsymbol{\tau} \cdot \mathbf{n})$ must be added to the surface integral term. The body force, \mathbf{f} , is force per unit mass.

Roadmap - Integral Relations



Chapter 2.5

Energy Equation



Energy Equation

Conservation of energy:

$$\underbrace{\frac{d}{dt} \iiint_{\Omega} \rho e_o d\mathcal{V}}_{\text{rate of change of total internal energy in } \Omega} + \underbrace{\iint_{\partial\Omega} [\rho e_o (\mathbf{v} \cdot \mathbf{n}) + p \mathbf{v} \cdot \mathbf{n}] dS}_{\text{net flow of total internal energy out from } \Omega \text{ plus work due to surface pressure on } \partial\Omega} = \underbrace{\iiint_{\Omega} \rho \mathbf{f} \cdot \mathbf{v} d\mathcal{V}}_{\text{work due to forces inside } \Omega}$$

where

$$\rho e_o = \rho \left(e + \frac{1}{2} \mathbf{v} \cdot \mathbf{v} \right) = \rho \left(e + \frac{1}{2} v^2 \right)$$

is the total internal energy

Energy Equation

The surface integral term may be rewritten as follows:

$$\oiint_{\partial\Omega} \left[\rho \left(e + \frac{1}{2}v^2 \right) (\mathbf{v} \cdot \mathbf{n}) + \rho \mathbf{v} \cdot \mathbf{n} \right] dS$$

$$\Leftrightarrow$$

$$\oiint_{\partial\Omega} \left[\rho \left(e + \frac{p}{\rho} + \frac{1}{2}v^2 \right) (\mathbf{v} \cdot \mathbf{n}) \right] dS$$

$$\Leftrightarrow$$

$$\oiint_{\partial\Omega} \left[\rho \left(h + \frac{1}{2}v^2 \right) (\mathbf{v} \cdot \mathbf{n}) \right] dS$$

Energy Equation

Introducing total enthalpy

$$h_o = h + \frac{1}{2}v^2$$

we get

$$\frac{d}{dt} \iiint_{\Omega} \rho e_o d\mathcal{V} + \oint_{\partial\Omega} [\rho h_o \mathbf{v} \cdot \mathbf{n}] dS = \iiint_{\Omega} \rho \mathbf{f} \cdot \mathbf{v} d\mathcal{V}$$

Energy Equation

Note 1: to include friction work on $\partial\Omega$, the energy equation is extended as

$$\frac{d}{dt} \iiint_{\Omega} \rho e_o d\mathcal{V} + \iint_{\partial\Omega} [\rho h_o \mathbf{v} \cdot \mathbf{n} - (\boldsymbol{\tau} \cdot \mathbf{n}) \cdot \mathbf{v}] dS = \iiint_{\Omega} \rho \mathbf{f} \cdot \mathbf{v} d\mathcal{V}$$

Note 2: to include heat transfer on $\partial\Omega$, the energy equation is further extended

$$\frac{d}{dt} \iiint_{\Omega} \rho e_o d\mathcal{V} + \iint_{\partial\Omega} [\rho h_o \mathbf{v} \cdot \mathbf{n} - (\boldsymbol{\tau} \cdot \mathbf{n}) \cdot \mathbf{v} + \mathbf{q} \cdot \mathbf{n}] dS = \iiint_{\Omega} \rho \mathbf{f} \cdot \mathbf{v} d\mathcal{V}$$

where \mathbf{q} is the heat flux vector

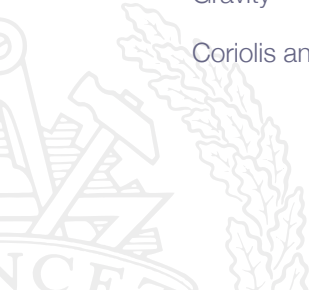
Energy Equation

Note 3: the force \mathbf{f} inside Ω may be a distributed body force field

Examples:

Gravity

Coriolis and centrifugal acceleration terms in a rotating frame of reference



Energy Equation

Note 4: there may be objects inside Ω which we choose to represent as sources of momentum and energy.

For example, there may be a solid object inside Ω which acts on the fluid with a force \mathbf{F} and performs work \dot{W} on the fluid

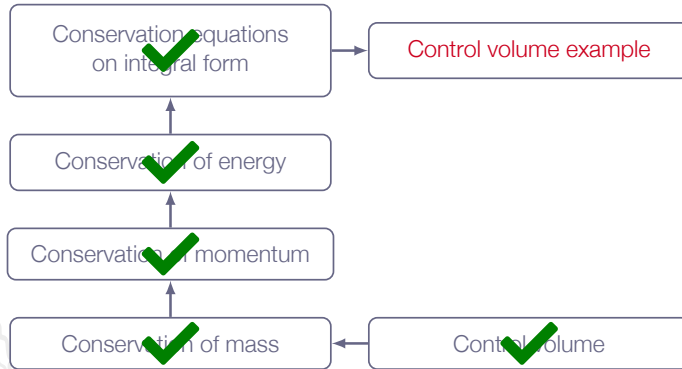
Momentum equation:

$$\frac{d}{dt} \iiint_{\Omega} \rho \mathbf{v} d\mathcal{V} + \iint_{\partial\Omega} [\rho(\mathbf{v} \cdot \mathbf{n})\mathbf{v} + p\mathbf{n}] dS = \iiint_{\Omega} \rho \mathbf{f} d\mathcal{V} + \mathbf{F}$$

Energy equation:

$$\frac{d}{dt} \iiint_{\Omega} \rho e_o d\mathcal{V} + \iint_{\partial\Omega} [\rho h_o \mathbf{v} \cdot \mathbf{n}] dS = \iiint_{\Omega} \rho \mathbf{f} \cdot \mathbf{v} d\mathcal{V} + \dot{W}$$

Roadmap - Integral Relations



Integral Equations - Applications

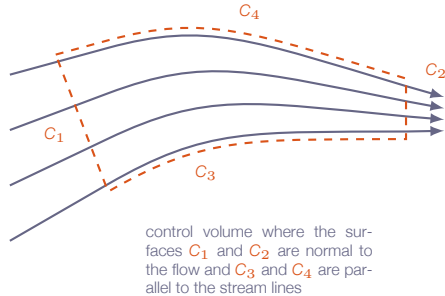
How can we use control volume formulations of conservation laws?

- ▶ Let $\Omega \rightarrow 0$: In the limit of vanishing volume the control volume formulations give the **P**artial **D**ifferential **E**quations (PDE:s) for mass, momentum and energy conservation (see Chapter 6)
- ▶ Apply in a "smart" way \Rightarrow Analysis tool for many practical problems involving compressible flow (see Chapter 2, Section 2.8)



Integral Equations - Applications

Example: Steady-state adiabatic inviscid flow



Integral Equations - Applications

Conservation of mass:

$$\underbrace{\frac{d}{dt} \iiint_{\Omega} \rho d\mathcal{V}}_{=0} + \underbrace{\oiint_{\partial\Omega} \rho \mathbf{v} \cdot \mathbf{n} dS}_{-\rho_1 v_1 A_1 + \rho_2 v_2 A_2} = 0$$

Conservation of energy:

$$\underbrace{\frac{d}{dt} \iiint_{\Omega} \rho e_o d\mathcal{V}}_{=0} + \underbrace{\oiint_{\partial\Omega} [\rho h_o \mathbf{v} \cdot \mathbf{n}] dS}_{-\rho_1 h_{o1} v_1 A_1 + \rho_2 h_{o2} v_2 A_2} = 0$$

Integral Equations - Applications

Conservation of mass:

$$\rho_1 v_1 A_1 = \rho_2 v_2 A_2$$

Conservation of energy:

$$\rho_1 h_{o1} v_1 A_1 = \rho_2 h_{o2} v_2 A_2$$

$$\Leftrightarrow$$

$$h_{o1} = h_{o2}$$

Total enthalpy h_o is conserved along streamlines in steady-state adiabatic inviscid flow

Roadmap - Integral Relations

